

Parallel spinors, Calabi–Yau manifolds and spinors

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The dominant energy condition

Let h be a Lorentzian metric on N

Energy-momentum tensor or Einstein tensor

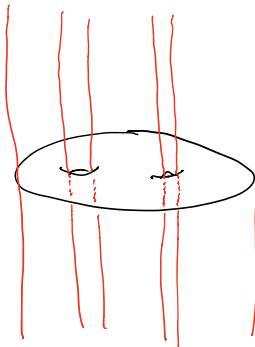
$$T^h := \text{Ric}^h - \frac{1}{2} \text{scal}^h h$$

We say that h satisfies the **dominant energy condition** in $x \in N$ if for all causal future oriented vectors $X, Y \in T_x N$:

$$T(X, Y) \geq 0. \quad (\text{DEC})$$

Physical interpretation (Einstein equation):

Non-negative mass density of matter fields.



Lorentz manifold
 (N, h) diffeo
to $M \times (a, b)$

compact space-
like hypersurface
 (M, g)

DEC on spacelike hypersurfaces

If M is a space-like hypersurface with induced metric g , and future-oriented unit normal, then we define:

Energy density $\rho := T^h(\nu, \nu) = \frac{1}{2} (\text{scal } g + (\text{tr } W)^2 - \text{tr}(W^2))$

Momentum density $j := T^h(\nu, \cdot)|_{T_x M} = \text{div } W - d \text{tr } W$

DEC for h implies $\rho \geq |j|$.

Definition

Let g be a Riemannian metric and W a g -symmetric endomorphism section. We say that (g, W) satisfies

- ▶ the **dominant energy condition** if $\rho \geq |j|$ (DEC)

$$\mathcal{I}^{\geq}(M) := \{(g, W) \text{ satisfying (DEC)}\}.$$

- ▶ the **strict dominant energy condition** if $\rho > |j|$ (DEC_>)

$$\mathcal{I}^{>}(M) := \{(g, W) \text{ satisfying (DEC}_{>})\}.$$

The inclusion $\mathcal{R}^{\geq}(M) \rightarrow \mathcal{I}^{\geq}(M)$

$$\mathcal{R}(M) \hookrightarrow \mathcal{I}(M), g \mapsto (g, 0)$$

$$\mathcal{R}^{\geq}(M) := \{g \in \mathcal{R} \mid \text{scal } g \geq 0\} = \mathcal{R}(M) \cap \mathcal{I}^{\geq}(M)$$

$$\mathcal{R}^{>}(M) := \{g \in \mathcal{R} \mid \text{scal } g > 0\} = \mathcal{R}(M) \cap \mathcal{I}^{>}(M)$$

Work by Jonathan Glöckle

A lot is known about $\mathcal{I}^{>}(M)$.

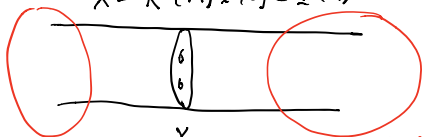
In particular, we have $(g, \lambda g) \in \mathcal{I}^{>}(M)$ if

- ▶ $g \in \mathcal{R}^{>}(M)$ and $\lambda \in \mathbb{R}$, or
- ▶ $g \in \mathcal{R}^{\geq}(M)$ and $\lambda \in \mathbb{R} \setminus \{0\}$, or
- ▶ $g \in \mathcal{R}(M)$ and $|\lambda| \gg 0$.

We get a map $\text{Susp}(\mathcal{R}^{>}(M)) \rightarrow \mathcal{I}^{>}(M)$.

$$\text{Susp}(\mathcal{R}^{>}(M)) = \left(\mathcal{R}^{>}(M) \times [-1, 1] / M \times \{-1\} \right) / M \times \{1\}.$$

$$X = \mathbb{R}^n(M) \times \{0\} \subset \mathbb{I}^n(M)$$



contractible
in $\mathbb{I}^n(M)$



contractible
in $\mathbb{I}^n(M)$

\rightsquigarrow



Susp(X)

The Lorentzian α -index

For any $\Psi : S^{k+1} \rightarrow \mathcal{I}^>(M)$ J. Glöckle constructs $\alpha_{\text{Lor}}(\Psi) \in \text{KO}^{-n-(k+1)-1,1}(\{*\}) \cong \text{KO}^{-n-k-1}(\{*\})$.

Theorem (J. Glöckle 2019)

The diagram

$$\begin{array}{ccccc} \pi_k(\mathcal{R}^>(M)) & \xrightarrow{\text{Susp}} & \pi_{k+1}(\text{Susp}(\mathcal{R}^>(M))) & \longrightarrow & \pi_{k+1}(\mathcal{I}^>(M)) \\ & \searrow \alpha_{\text{Riem}} & & \swarrow \alpha_{\text{Lor}} & \\ & & \text{KO}^{-n-k-1}(\{*\}) & & \end{array}$$

commutes.

Key technique in Glöckle's article: The Dirac-Witten operator

Literature: Witten 1981, Parker-Taubes, Hijazi-Zhang, . . . , Glöckle 2019.

Restrict the spinor bundle ΣN from (N, h) to (M, g) .

As spinor module $\Sigma N|_M$ is one or two copies of ΣM .

However:

scalar product $\langle\langle \cdot, \cdot \rangle\rangle$ on ΣN is indefinit (splitt signature),

scalar product $\langle \cdot, \cdot \rangle$ on ΣM positiv definit.

They are related by

$$\langle \varphi, \psi \rangle = \langle\langle \nu \cdot \varphi, \psi \rangle\rangle.$$

The connections differ:

$$\nabla_X^N \varphi = \nabla_X^M \varphi - \frac{1}{2} \nu \cdot W(X) \cdot \varphi$$

Dirac-Witten-Operator

$$D^{(g,W)} \varphi = \sum_{j=1}^n e_j \cdot \nabla_{e_j}^N \varphi$$

where (e_1, \dots, e_n) is a locally defined orthonormal frame of TM .



$D^{(g,W)}$ is self-adjoint and Fredholm.

Schrödinger-Lichnerowicz formula:

$$\left(D^{(g,W)}\right)^2 = (\nabla^N)^* \nabla^N + \frac{1}{2}(\rho - \nu \cdot j^\sharp \cdot),$$

The $*$ is taken on M with respect to $\langle \cdot, \cdot \rangle$.

Recall:

Energy density $\rho := T^h(\nu, \nu) = \frac{1}{2} (\text{scal } g + (\text{tr } W)^2 - \text{tr}(W^2))$

Momentum density $j := T^h(\nu, \cdot)|_{T_x M} = \text{div } W - d \text{tr } W$

DEC for h implies $\rho \geq |j|$.

This implies that $D^{(g,W)}$ is invertible if $(g, W) \in \mathcal{I}^>(M)$.

As a consequence Glöckle can use index theoretical methods.

Our goal

Question

Want to understand $\mathcal{I}^{\geq}(M)$.

Does this has similar properties as $\mathcal{I}^>(M)$?

Compare: Any non-trivial element in $\pi_k(\mathcal{R}^>(M))$, detected by index theory remains non-trivial in $\pi_k(\mathcal{R}^{\geq}(M))$.

In fact, suppose we have a map, $k > 0$,

$$g : S^k \rightarrow \mathcal{R}^{\geq}(M), \quad \alpha \mapsto g_{\alpha},$$

and assume $g_{\beta} \in \mathcal{R}^>(M)$ for some $\beta \in S^k$.

Then Schick–Wraith showed that $D^{g_{\alpha}}$ is invertible for all $\alpha \in S^k$.

Important ingredients:

- ▶ if $\varphi \in \ker D^{g_{\alpha}}$, then $\nabla\varphi = 0$ (“the kernel case”)
- ▶ Rigidity for metrics with parallel spinors (McKenzie Wang / Dai–Wang–Wei)

Alternative presentation of our question

In the following diagram we assume $k \geq 1$ and that the base point is g_0 resp. $(g_0, 0)$ where g_0 has positive scalar curvature.

$$\begin{array}{ccc} \pi_k(\mathcal{R}^>(M)) & \longrightarrow & \pi_{k+1}(\mathcal{I}^>(M)) \\ \downarrow & & \downarrow \\ \pi_k(\mathcal{R}^{\geq}(M)) & \longrightarrow & \pi_{k+1}(\mathcal{I}^{\geq}(M)) \end{array}$$

Index theoretically determined non-trivial homotopy groups survive in upper right and in lower left corner.

What about the lower right corner?

Analogous results for the Dirac–Witten operator?

Proposition (Ammann, Glöckle)

Assume that M is a connected closed spin manifold and $(g, W) \in \mathcal{I}^{\geq}(M)$.

We assume that $\varphi \in \ker D^{(g,W)} \setminus \{0\}$.

Then g, W, φ provides initial data for a Lorentzian manifold with a parallel spinor.

Example

If $(M, g, W) \subset (N, h)$ Lorentz manifold, $\Phi \neq 0$ a parallel spinor. Then $\nabla^N \Phi|_M = 0$, and thus $\Phi|_M \in \ker D^{(g,W)}$.

Moreover, we then have $\text{Ric} = f\alpha \otimes \alpha$ for some lightlike α . Thus $\text{scal}^h = 0$ and $T = f\alpha \otimes \alpha$.

$$f \leq 0 \Leftrightarrow (g, W) \in \mathcal{I}^{\geq}(M).$$

Dirac currents

The Dirac current of a Lorentzian manifold (N, h) is the vector field V_φ with

$$h(X, V_\varphi) = -\langle\langle X \cdot \varphi, \varphi \rangle\rangle \quad \forall X \in TN$$

Then $h(V_\varphi, V_\varphi) \leq 0$, i.e. V_φ is causal.

If $V_\varphi(p)$ is lightlike, then $V_\varphi(p) \cdot \varphi(p) = 0$.

The Dirac current of a Riemannian manifold (M, g) is the vector field U_φ with

$$g(X, U_\varphi) = i\langle X \bullet \varphi, \varphi \rangle \quad \forall X \in TN$$

Unfortunately on $\Sigma N|_M$ two different Clifford multiplications are used in the literature:

· is given as the pullback of the Clifford multiplication on N .

$$X \bullet \varphi = i\nu \cdot X \cdot \varphi.$$

Then $V_\varphi|_M = -U_\varphi + u_\varphi \nu$ for some $u \in C^\infty(M)$.



Spacelike hypersurfaces

Work by H. Baum, T. Leistner, A. Lischewski

If (N, h) is a **Lorentzian** manifold with a parallel spinor φ .

As φ is parallel, V_φ is a parallel vector field.

Assuming M connected V_φ is either timelike or lightlike everywhere.

If V_φ is timelike, we locally have $N = M_0 \times \mathbb{R}$, $h = g - dt^2$.

The easier case.

Lightlike case

We assume V_φ is lightlike.

$$\begin{aligned}\nabla^N \varphi &= 0, \\ V_\varphi \cdot \varphi &= 0\end{aligned}$$

If we “restrict” φ to M , these equations imply the constraint equations

$$\begin{aligned}\nabla_X^M \varphi &= \frac{i}{2} W(X) \bullet \varphi, & \forall X \in TM, \\ U_\varphi \bullet \varphi &= iu_\varphi \varphi,\end{aligned}\tag{CE}$$

The Cauchy problem for parallel spinors

Conversely, if we have a Riemannian manifold (M, g) with a non-trivial solution of

$$\begin{aligned}\nabla_X^M \varphi &= \frac{i}{2} W(X) \bullet \varphi, & \forall X \in TM, \\ U_\varphi \bullet \varphi &= iu_\varphi \varphi,\end{aligned}\tag{CE}$$

then it extends to a Lorentzian metric on $M \times (-\epsilon, \epsilon)$ with a parallel spinor φ with V_φ lightlike.

Again: work by H. Baum, T. Leistner, A. Lischewski
Simplified by Julian Seipel (Master thesis, Regensburg),
following ideas by P. Chrusciel

Remark. In my last talk in Freiburg I explained:

Ammann–Kröncke–Müller proved:

For any family of metrics k_τ , $a < \tau < b$ on Q with a parallel spinors, we obtain a solutions to (CE) on

$$(M = Q \times (a, b), g = f_\tau^*(k_\tau) + d\tau^2). \quad (*)$$

Leistner–Lischewski proved:

For any solution to (CE), M is locally isometric to a generalized version of (*).

Thus solutions to (CE) on m -dimensional manifolds are tightly related to families of metrics with parallel spinor on $(m - 1)$ -dimensional manifolds.

Some steps in the proof

Assume (DEC) and $\varphi \in \ker D^{(g,W)}$.

The Schrödinger-Lichnerowicz equation implies

Schrödinger-Lichnerowicz formula:

$$\begin{aligned} 0 &= \int_M \langle D^{(g,W)}\varphi, D^{(g,W)}\varphi \rangle d\mu^M \\ &= \underbrace{\int_M \langle \nabla^N \varphi, \nabla^N \varphi \rangle d\mu^M}_{\geq 0} + \frac{1}{2} \underbrace{\int_M \langle (\rho - \nu \cdot j^\sharp) \varphi, \varphi \rangle d\mu^M}_{\geq 0}, \end{aligned}$$

This implies $\nabla^N \varphi = 0$, $\rho\varphi = \nu \cdot j^\sharp \cdot \varphi = -i j^\sharp \bullet \varphi$, and $\rho = |j|$.

The Lorentzian Dirac current V_φ is well-defined and ∇^N -parallel along M .

Thus V_φ is everywhere lightlike or everywhere timelike.

The case $\rho \neq 0$

In this case a calculation shows on $\{x \in M \mid \rho(x) \neq 0\}$:

$$U_\varphi = \frac{j^\sharp}{\rho} \|\varphi\|^2.$$

The definition of V_φ implies

$$V_\varphi = -U_\varphi + \|\varphi\|^2 \nu,$$

and thus V_φ is lightlike. This implies (CE).

The case $\rho \equiv 0$

Now $\rho = 0$ and $j = 0$. This means that the constraint equations for vacuum Einstein equation are satisfied.

Thus we can choose N to be Ricci-flat (Choquet-Bruhat et al.). We then can extend φ to a parallel section of ΣN .

Note: In this case φ might be spacelike or timelike.

Summary and outlook

Riemannian	Lorentzian IDS
scal ≥ 0	DEC
scal > 0	DEC+
Dirac operator	Dirac–Witten operator
$\pi_k(\{\text{scal} > 0\}) \xrightarrow{\alpha} KO_{m+k+1}(p)$	$\pi_{k+1}(\{\text{scal} > 0\}) \xrightarrow{\alpha} KO_{m+k+1}(p)$
no scal > 0 metric	no path from $(g, -Ng)$ to (g, Ng)
parallel spinor	initial data for Lorentzian manifolds with parallel spinors
smooth finite dim. moduli space	???
Wang rigidity	???
$\pi_k\{\text{scal} \geq 0\} \xrightarrow{\alpha} KO_{m+k+1}(p)$???